# A COMPARATIVE STUDY OF SSN AND DST FOR SOLAR CYCLE 23 AND 24

Rajesh Mathpal<sup>1</sup>, Lalan Prasad<sup>2</sup>, 1Department of Physics, M.B. Govt. P.G. College, Haldwani (Nainital) 263 139, India 2Department of Physics, Govt P.G. College, Berinag (Pithoraghar) 263 125, India

#### **ABSTRACT-**

In the present study we have investigated the impact of sunspot numbers on geomagnetic stroms. In our study Disturbance storm time (Dst) index is taken as a measure of Geomagnetic Storm (GS). We have investigated SSN with Dst ( $\leq$  -50 nT) during the time interval 1997-2016. The comparative study of dst and SSN shows V shaped structure. The study shows symmetric and asymmetric structures. This shows the variation in geomagnetic field is greatly affected by various phase of solar cycle. The phase relationship caused due to differential rotations of Sun causes change in ring current shielding the earth. This causes change in geomagnetic activity. This study is important for studying the space weather phenomenon.

Keywords: GS - Geomagnetic storm, Dst - Disturbance storm-time, CME - Coronal mass ejection, ICME - Interplanetary coronal mass ejection, SSN – Sunspot Number.

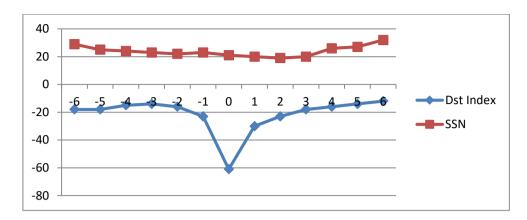
#### INTRODUCTION-

The Dst index is a measure in the context of space weather. It gives information about strength of ring current around earth caused by solar protons and electrons. The ring current around earth produces magnetic field that is directly opposite Earth's magnetic field i.e. if the difference between solar electron and protons gets higher then Earth's magnetic field becomes weaker. During a geomagnetic storm, the number of particles in the ring current increases which causes a decrease in geomagnetic field. During a geomagnetic storm, the number of particles in the ring current increases which causes a decrease in geomagnetic field. The solar wind is highly correlated to Dst in mid of solar cycle-23 which represents the maximum solar activity years (Mathpal et al., 2016). The relation varies during solar cycle where the GS occurs within 5 days after the onset of coronal mass ejection and variation in geomagnetic disturbance generally follows the phase of the solar cycle (Prasad et al., 2013b).

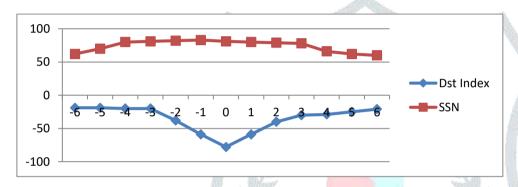
The relation varies during solar cycle where the GS occurs within 5 days after the onset of coronal mass ejection and variation in geomagnetic disturbance generally follows the phase of the solar cycle (Prasad et al., 2013b). The Interplanetary magnetic field Bz component is generated as a result of solar wind turbulence and wave generation (Velli and Birn 1999). Aschwanden (2006) stated the modern definition of CMEs as "a dynamically evolving plasma structure propagating outward from the sun into interplanetary space, carrying frozen-in magnetic flux and expanding in size". Generally when CME propagates, it enhances the speed of solar wind. The solar wind is an important parameter for occurrence of geomagnetic disturbance. When CME from Sun propagate in interplanetary space it creates disturbances in solar wind. Effect of Solar Flare and Coronal Mass Ejection in our Earth studied (Prasad L et al. 2014).

Inter planetary Coronal Mass Ejection (ICME) when arrive near the earth it contains solar wind plasma ahead of high speed solar wind, this compressed solar wind plasma exerts a dynamic pressure on Earth's magnetosphere (Prasad et al.2013). Many investigators have studied the association of geomagnetic activity to magnetic clouds and other IMF features (Farrugia, Burlaga, and Lepping, 1997; Tsurutani and Gonzalez, 1997).

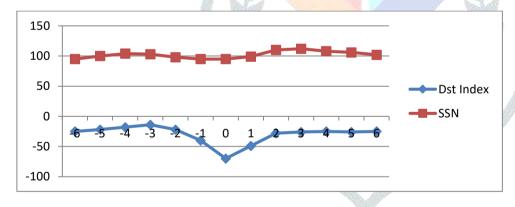
# DATA COLLECTION AND ANALYSIS-



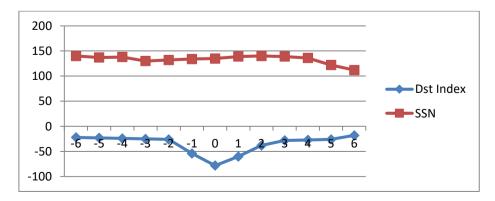
**YEAR: 1997** 



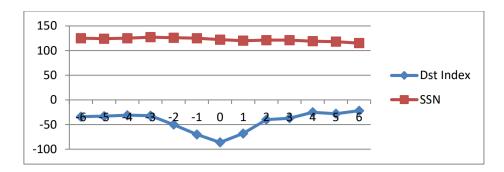
**YEAR: 1998** 



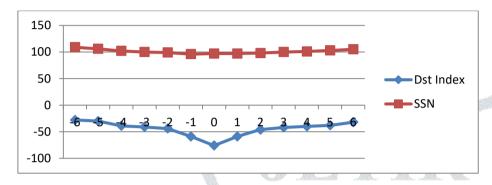
**YEAR: 1999** 



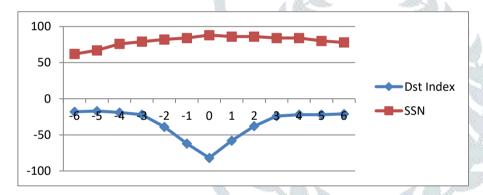
**YEAR: 2000** 



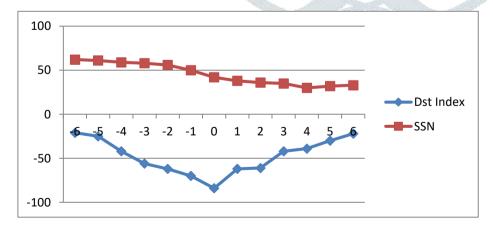
**YEAR: 2001** 



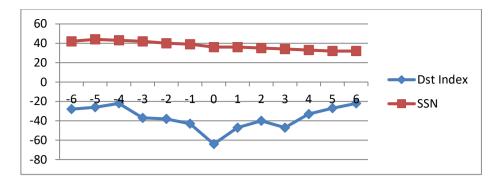
**YEAR: 2002** 



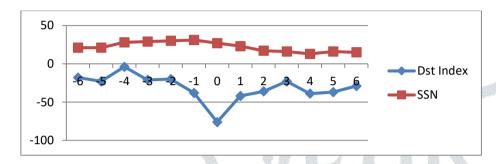
**YEAR: 2003** 



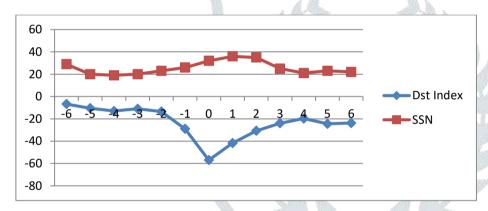
**YEAR: 2004** 



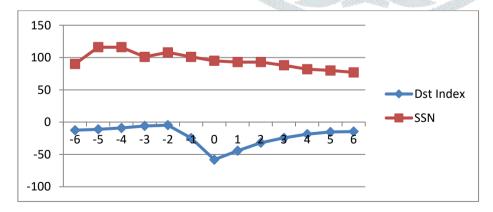
**YEAR: 2005** 



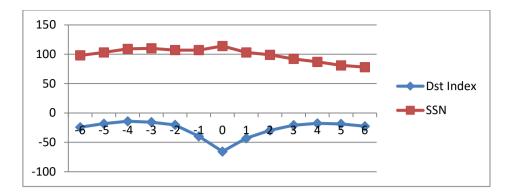
**YEAR: 2006** 



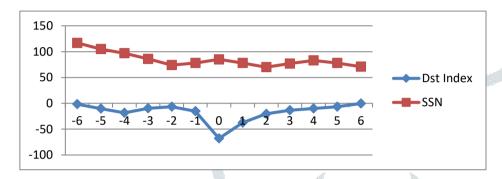
**YEAR: 2010** 



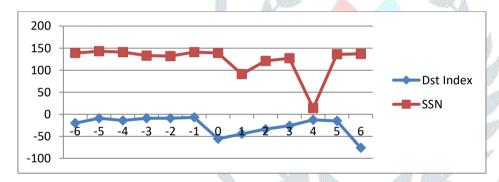
**YEAR: 2011** 



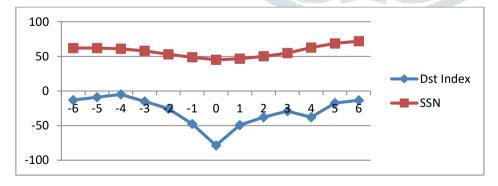
**YEAR: 2012** 



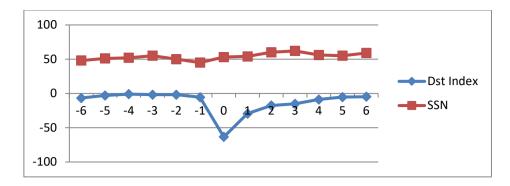
**YEAR: 2013** 



**YEAR: 2014** 

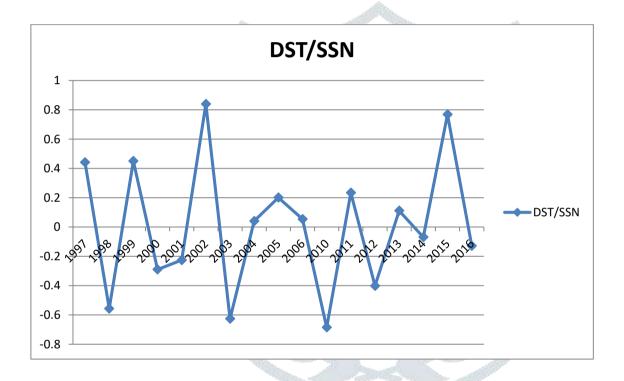


**YEAR: 2015** 



**YEAR: 2016** 

### **CORRELATION GRAPH -**



S. No.	Parameters	Solar cycle	Correlation coefficient
1	Dst index and SSN	23	0.032703
2	Dst index and SSN	24	-0.02461

# **Results** -

- 1) For 2002 dst is highly correlated with SSN. This indicates high SSN for this year weakened the earth's magnetic field which causes high ring current.
- 2) Similar peak is obtained for year 2015.
- 3) The correlation becomes weak at the end of 23 solar cycles, disappears for 2007,2008 and 2009.
- 4) For solar cycle 24 correlation rise from low to high.

- 5) For years 1998, 2003 and 2010 it is highly negatively correlated.
- 6) For years 1997 and 1998 there is good correlation.

**Conclusion** – The comparative study of dst and SSN shows V shaped structure. The study shows symmetric and asymmetric structures. This shows the variation in geomagnetic field is greatly affected by various phase of solar cycle. The phase relationship caused due to differential rotations of Sun causes change in ring current shielding the earth. This causes change in geomagnetic activity. This study is important for studying the space weather phenomenon. This study also suggests that solar storms are influenced by SSN.

ACKNOWLEDGEMENTS – The authors are thankful to IUCAA Pune for content page service and also thankful to ARIES, Nainital for providing library and computing facilities. We are also thankful to website http://omniweb.gsfc.nasa.gov/cgi/nx1.cgi. for the data collection.

#### REFRENCES-

Badruddin B. 2002, Solar Physics, 165, 195206.

Badruddin B. and Yadav, R. S. 1982, Indian J. Physics, 68, 588.

Gonzalez, W.D., Joselyn, J.A., Kamide, Y., Kroehl, H.W., Rostoker, G., Tsurutani, B.T., Vasyliunas, V.M.: 1994, J. Geophys. Res. 99, 5771, 1994.

Gopalswamy, N., Yashiro, S., Akiyama, S., 2007, J. Geophys. Res., 112, A06112. Hotton, C. J. 1980, Solar Phys., 66, 159.

Howard, R.A., Michels, D.J., Jr. Sheeley, N.R., Koomen, M.J., 1982, Astrophys. J. Lett., 263, L101.

Kharayat, H., Prasad, L., Mathpal, R., Garia, S., Bhatt, B.: 2016, Solar Phys. 291, 603. Lockwood, J.A. and Webber, W.R., 1992, J. Geophys. Res., 97: 8221-8230.

Mathpal, R., Prasad, L., Kharayat, H., Bhoj, C., Mathpal, C. & Pokharia, M. 2016, International Journal of Physics and Astronomy, 28 (1), 1184-1190.

Parsai, N. and Singh, N. 2014, Int. J. Theoretical & Applied Sci. 6(2), 10-12.

Prasad, L., Joshi, V. K., 2008, *Physics Education*, 25, 267.

Prasad, L., Garia, S., Bhatt, B., 2013, Int. J. Phys. Appl., 5(2), 77-81.

Shrivastava, P.K. and Jaiswal, K. L. 2003, Solar Physics, 214, 195-200.

Shrivastava P.K., Jothe M.K. and Singh M., 2011, Solar Physics, 269, 401-410.